

# Environmental Fluid Dynamics: Radiation Model

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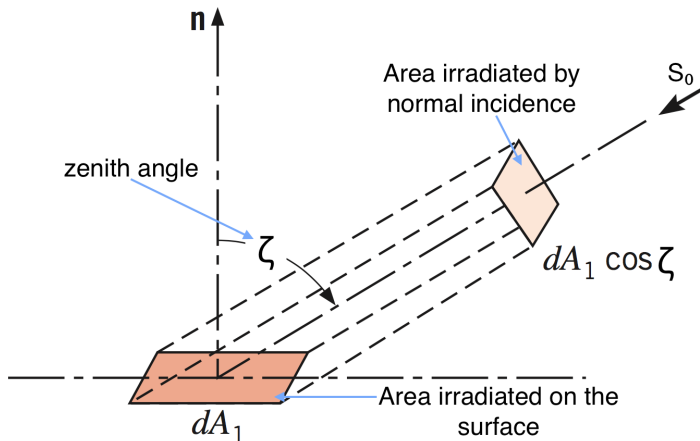


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# A Simple Model for the Computation of Radiation on a Slope

# Radiation Model: Overview



- Flux density on a horizontal surface

$$S = S_0 \cos \zeta \rightarrow \text{Lambert's Cosine Law}$$

where  $S_0 = 1367 \text{ W m}^{-2}$  is the solar constant



**We want to consider a more general case of radiation on a slope of arbitrary:**

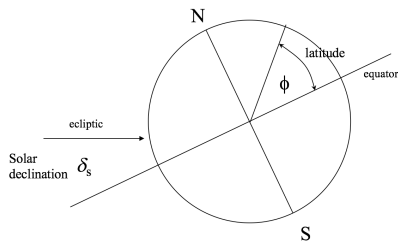
- angle
- orientation
- time of day
- location on Earth
- time of year

**Let's build a model:**

- Consider seasonal effects (Earth's orbit)
- Daily effects (local sun angle and azimuth)
- Slope (mountains, walls, etc)
- Atmospheric composition



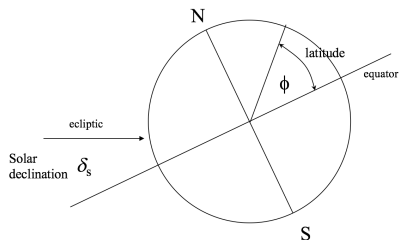
# Radiation Model: Seasonal Effects



- $\phi$  - latitude
- $\phi_r$  - tilt of the Earth's axis relative to the ecliptic (orbital plane of the Earth around the sun)
- $\phi_r = 23.45^\circ = 0.409$  rad corresponds to the latitude of the tropics (Capricorn and Cancer)



# Radiation Model: Seasonal Effects



## Solar declination $\delta_s$

- Angle between the ecliptic and the equator as the Earth rotates around the sun
- Summer and Winter

$$\underbrace{-23.45^\circ}_{\text{winter solstice}} \lesssim \delta_s \lesssim \underbrace{23.45^\circ}_{\text{summer solstice}}$$

- Spring and Autumn  
 $\delta_s = 0$  means sun is directly overhead at equator:  
equinox  
vernal (Mar 19-21)  
autumnal (Sept 22-24)



## Solar declination $\delta_s$

- For a circular orbit:

$$\delta_s = \phi_r \cos \left[ \frac{C(d - d_r)}{d_y} \right]$$

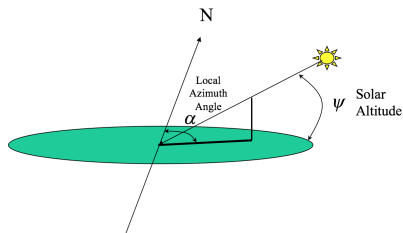
where

- $C = 360^\circ$  or  $2\pi$  rad
- $d =$  day of the year (Julian date 1-365, or 366)
- $d_r =$  summer solstice (June 22, 173 JD, 20-22)
- $d_y = 365$  or 366 days
- $\phi_r =$  tilt of Earth's axis relative to the ecliptic ( $23.45^\circ$ )





# Radiation Model: Daily Effects



- $\psi$  - local sun elevation angle (“solar altitude”)
- $\alpha$  - local azimuth angle ( $> 0$  clockwise)
- $\phi$  - latitude ( $> 0$  N hemisphere)
- $\lambda_e$  - longitude ( $> 0$  W of prime meridian)



## Spherical relationships for local elevation angle

$$\sin \psi = \sin \phi \sin \delta_s - \cos \phi \cos \delta_s \cos \underbrace{\left[ \frac{Ct_{\text{UTC}}}{t_d} - \lambda_e \right]}_B$$

where

- $C = 360^\circ$  or  $2\pi$  rad
- $t_{\text{UTC}} = \text{local time} + \text{hours to UTC}$
- $t_d = 24$  hours
- $\lambda_e = \text{longitude}$  ( $> 0$  west from prime meridian)



## Spherical relationships for local azimuth angle

$$\cos \alpha = \frac{\sin \delta_s - \sin \phi \cos \zeta}{\cos \phi \sin \zeta}$$

where

- the zenith angle  $\zeta$  is given by

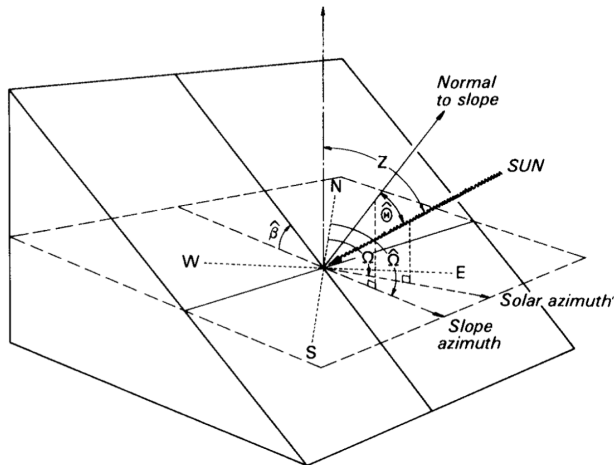
$$\begin{aligned}\zeta &= 90^\circ - \psi \\ &= \frac{\pi}{2} - \psi \text{ (radians)}\end{aligned}$$

- Correction in afternoon for sun setting in west is

$$\alpha' = C - \alpha$$



# Radiation Model: Slope Effects



- $Z$  - Zenith angle
- $\hat{\beta}$  - Slope angle
- $\hat{\Omega}$  - Solar azimuth angle
- $\hat{\Omega}$  - Slope azimuth angle
- $\hat{\theta}$  - Angle of incidence  
(between Sun and the  
normal to the slope)



## Spherical relationships for slope

$$\cos \hat{\theta} = \cos \hat{\beta} \cos \zeta + \sin \hat{\beta} \sin \zeta \cos(\alpha - \hat{\Omega})$$

Recall that for a flat surface

$$S = S_0 \cos \zeta$$

While for a sloping surface (absent scattering/absorption)

$$S = S_0 \cos \hat{\theta}$$



## Net shortwave radiation

$$R_s = R_{s\downarrow}(1 - a)$$

if  $|R_L| \ll R_s$  (true under clear skies during the day), then

$$R_N = R_{s\downarrow}(1 - a)$$

In reality, we must also consider atmospheric transmissivity

$$\begin{aligned} R_{s\downarrow} &= S_0 T_r \sin \phi \\ &= S_0 T_r \cos \zeta \end{aligned}$$

where

- $T_r$  = transmissivity (net sky transmissivity)
- $S_0$  = solar constant  $\simeq 1367 \text{ W m}^{-2}$



**Transmissivity** Depends on

- path length through atmosphere
- atmospheric absorption
- cloudiness

Simple model

$$T_r = (0.6 + 0.2 \sin \psi)(1 - 0.4\sigma_H)(1 - 0.7\sigma_M)(1 - 0.4\sigma_L)$$

where  $\sigma_H$ ,  $\sigma_M$ , and  $\sigma_L$  are cloud cover fractions for high-, mid-, and low-level clouds, respectively ( $0 \leq \sigma_i \leq 1$ )



## Net longwave radiation

$$R_L = R_{L\downarrow} + R_{L\uparrow}$$

where

$$R_{L\uparrow} = -\epsilon\sigma T^4$$

However, it is harder to model  $R_{L\downarrow}$ , so we will model the net longwave radiation  $R_L$

$$R_L = b(1 - 0.1\sigma_H - 0.3\sigma_M - 0.6\sigma_L)$$

where  $b = -98.5 \text{ W m}^{-2}$





## Net radiation model

$$\begin{aligned}R_N &= R_{S\downarrow} + R_{S\uparrow} + R_{L\downarrow} + R_{L\uparrow} \\ &= S_0 Tr \cos \zeta (1 - a) + R_L \\ &= R_L (R_L < 0)\end{aligned}$$

during day

during night



# Radiation Model: Net Radiation Model

If asked to model  $R_N$  for a specific location, on a specific day, at a specific time, with specific cloud cover and surface properties, follow these steps

- Compute  $\delta_s$
- Find solar elevation angle  $\psi$
- Solve for transmissivity  $T_r$
- Compute  $R_S$  and  $R_L$  contributions
- Compute  $R_N$

